Black hole thermodynamics in the presence of NLE fields

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BLACK HOLE THERMODYNAMICS

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0.
$$\kappa, \Phi, \ldots = \text{const. on } H$$

1.
$$\delta M = \frac{1}{8\pi} \kappa \delta A + \Phi_{\mathsf{H}} \delta Q + \dots$$

Smarr
$$M = \frac{1}{4\pi} \kappa A + \Phi_H Q + \dots$$

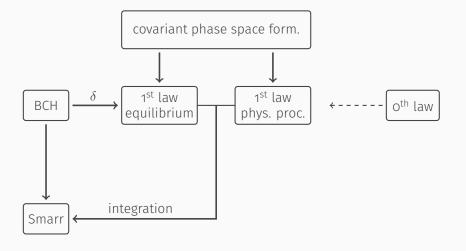
...and an auxiliary relation:

Bardeen-Carter-Hawking mass formula

$$M = \frac{1}{4\pi} \kappa \mathcal{A} + 2\Omega_{\mathsf{H}} J - 2 \int_{\Sigma} \left(\star T(\chi) - \frac{1}{2} T \star \chi \right)$$

...for a Killing horizon $H[\chi]$ generated by χ^a

RELATIONS AMONG FORMULAE



GOING BEYOND EINSTEIN AND MAXWELL

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- Maxwell's electrodynamics
 - → NONLINEAR ELECTRODYNAMICS

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 $\mathcal{F}\mathcal{G}$ -class $\mathscr{L}(\mathcal{F},\mathcal{G})$

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- ► NLE Lagrangians:

$$\begin{array}{ll} {\mathcal F}\text{-class} & {\mathcal L}({\mathcal F}) \\ \\ {\mathcal F}{\mathcal G}\text{-class} & {\mathcal L}({\mathcal F},{\mathcal G}) \end{array}$$

- ▶ auxiliary 2-form: $Z_{ab} := -4(\mathscr{L}_{\mathfrak{F}} F_{ab} + \mathscr{L}_{\mathfrak{G}} \star F_{ab})$
- generalized Maxwell's eqs.: $d\mathbf{F} = 0$, $d \star \mathbf{Z} = 0$



Max Born, Leopold Infeld, Werner K. Heisenberg, Hans H. Euler



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 \blacktriangleright Max Born (1933): an upper limit b for the field strength

$$\mathscr{L}^{(\mathsf{Born})} = b^2 \left(1 - \sqrt{1 + \frac{\mathcal{F}}{2b^2}} \right)$$

► Born-Infeld (1934)

$$\mathcal{L}^{(BI)} = b^2 \left(1 - \sqrt{1 + \frac{\mathcal{F}}{2b^2} - \frac{\mathcal{G}^2}{16b^4}} \right)$$

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► Euler-Heisenberg (1935): one-loop QED corrections to classical Maxwell

$$\mathcal{L}^{(EH)} = -\frac{1}{4} \mathcal{F} + \frac{\alpha^2}{360 m_e^4} \left(4\mathcal{F}^2 + 7\mathcal{G}^2 \right) + O(\alpha^3)$$

THIS TALK: RECENT RESULTS

BJS 2021

A. Bokulić, T. Jurić and I.S.:

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PRD 103 (2021) 124059 [2102.06213]

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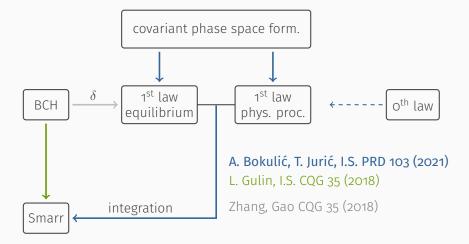
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GS 2018

L. Gulin and I.S.: **Generalizations of the Smarr formula for black holes with NLE fields** CQG **35** (2018) 025015 [1710.04660]

RECENT RESULTS



Zeroth law(s)

surface gravity $K^b \nabla_{\!b} K^a \stackrel{H}{=} \kappa K^a$ electric potential $E := -i_{\rm K} F = -{
m d}\Phi$ magnetic potential $H := i_{\rm K} \star Z = -{
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m d}\Psi$

$$\kappa, \Phi, \Psi \stackrel{H}{=} \text{const.}$$

proof via	κ	Φ,Ψ for Maxwell
Einstein's EOM	Bardeen et al. 1973	Carter 1973
bifurcation surface	Kay & Wald 1991	Gao 2003
Frobenius' theorem	Carter 1972	I.S. 2012, 2014

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 or $B_a = 0$

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$$E_a = 0$$
 or $B_a = 0$

$$D_a = K^b Z_{ab} = 0$$
 or $H_a = K^b \star Z_{ba} = 0$
if $(\mathcal{L}_{\mathcal{F}})^2 + (\mathcal{L}_{\mathcal{G}})^2 \neq 0$

First law

ightharpoonup variation of a field ϕ

$$\delta\phi(x):=\frac{\partial\phi(x;\lambda)}{\partial\lambda}\Big|_{\lambda=0}$$

 \blacktriangleright variation of a field ϕ

$$\delta\phi(x) := \frac{\partial\phi(x;\lambda)}{\partial\lambda}\Big|_{\lambda=0}$$

subtle issue: variation of coupling constants

$$\delta\beta_i := \frac{\partial\beta_i(\lambda)}{\partial\lambda}\Big|_{\lambda=0}$$

$$\delta \mathbf{L}[\phi; \beta] = \mathbf{E}[\phi; \beta] \,\delta \phi + \mathbf{\Lambda}^{i}[\phi; \beta] \,\delta \beta_{i} + \mathbf{d}\mathbf{\Theta}[\phi, \delta \phi; \beta]$$

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$$\mathbf{J}_{\xi} := \mathbf{\Theta}[\phi, \pounds_{\xi}\phi; \beta] - i_{\xi}\mathbf{L}[\phi; \beta]$$
$$d\mathbf{J}_{\xi} \approx 0 , \quad \mathbf{J}_{\xi} \approx d\mathbf{Q}_{\xi}$$

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$$\begin{aligned} \mathbf{J}_{\xi} &:= \mathbf{\Theta}[\phi, \pounds_{\xi} \phi; \beta] - i_{\xi} \mathbf{L}[\phi; \beta] \\ & \mathrm{d} \mathbf{J}_{\xi} \approx 0 \;, \quad \mathbf{J}_{\xi} \approx \mathrm{d} \mathbf{Q}_{\xi} \end{aligned}$$

$$\Omega_{\Sigma}[\phi, \delta\phi, \pounds_{\xi}\phi; \beta] = \int_{\Sigma} i_{\xi}(\mathbf{E}\,\delta\phi + \delta\mathbf{C}) +$$

$$+ \int_{\partial\Sigma} (\delta\mathbf{Q}_{\xi} - i_{\xi}\mathbf{\Theta}[\phi, \delta\phi; \beta]) - K_{\xi}^{i}(\beta)\,\delta\beta_{i}$$

$$K_{\xi}^{i}(\beta) := -\int_{\Sigma} i_{\xi}\mathbf{\Lambda}^{i}[\phi; \beta]$$

NLE FIRST LAW (BJS 2021)

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$$\delta M = \frac{\kappa}{8\pi} \delta \mathcal{A} + \Omega_{\mathsf{H}} \delta J + \Phi_{\mathsf{H}} \delta Q + \sum_{i} K_{\chi}^{i} \delta \beta_{i}$$

$$K_{\chi}^{i} = -\frac{1}{4\pi} \int_{\Sigma} \frac{\partial \mathcal{L}}{\partial \beta_{i}} \star \chi$$

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N.B.: Legendre transformation $M = \mathcal{E} + K_{\chi}^{i} \beta_{i}$

Smarr formula

GENERALIZED NLE SMARR (GS 2018)

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$$M = \frac{\kappa \mathcal{A}}{4\pi} + 2\Omega_{\mathsf{H}}J + \Phi_{\mathsf{H}}Q + \Psi_{\mathsf{H}}P + \Delta$$

$$\Delta = \frac{1}{2} \int_{\Sigma} T \star \chi$$

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...if
$$\mathcal{L} = \sigma^{-1} f(\sigma \mathcal{F}, \sigma \mathcal{G})$$
 then $\Delta = \sigma \mathcal{C}$

First Law \rightarrow Smarr formula

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• rescaled field configurations $(\lambda^2 g_{ab}, \lambda^{\nu} \mathbf{A})$

FIRST LAW → SMARR FORMULA

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- consistency: $\beta_i \to \lambda^{b_i} \beta_i$

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- rescaled field configurations $(\lambda^2 g_{ab}, \lambda^{\nu} \mathbf{A})$
- consistency: $\beta_i \to \lambda^{b_i} \beta_i$
- using $\delta \mathscr{Q} = \mathrm{d} \mathscr{Q}/\mathrm{d} \lambda \,|_{\lambda=1}$

$$M = \frac{\kappa}{4\pi} \mathcal{A} + 2\Omega_{\mathsf{H}} J + \Phi_{\mathsf{H}} Q + \sum_{i} b_{i} K_{\chi}^{i} \beta_{i}$$

crucial issue:

turning
$$T\star \chi$$
 into $d(\Phi\star \mathbf{Z})$, $d(\Phi\mathbf{F})$, $d(\Psi\mathbf{F})$ and $d(\Psi\star \mathbf{Z})$

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NLE with Maxwellian weak field limit, $\mathcal{L}_{\mathcal{F}}(0,0) = -1/4$ and $\mathcal{L}_{\mathcal{G}}(0,0) = 0 \Rightarrow$...Maxwell!

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- NLE with Maxwellian weak field limit, $\mathcal{L}_{\mathfrak{F}}(0,0)=-1/4$ and $\mathcal{L}_{\mathfrak{F}}(0,0)=0 \Rightarrow ...$ Maxwell!

Open questions

• generalize NLE Smarr with Λ , D=3, $D\geq 5$, etc.

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thermodynamic role of coupling parameters

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thermodynamic role of coupling parameters

▶ EiBI gravity w/ Maxwell ←→ GR w/ BI NLE

Thank you for the attention!

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