

# *Geometric Justification of the Fundamental Interaction Fields for the Classical Long-Range Forces*

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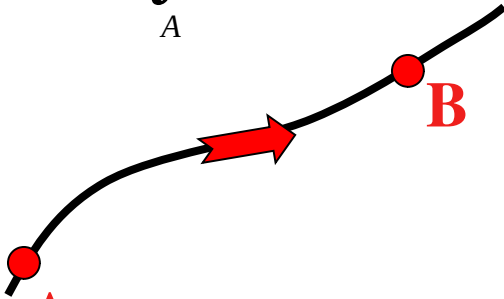
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# Relativistic Particle

$$L_1(x, v) = \sqrt{g_{\mu\nu} v^\mu v^\nu}, \quad L_2(x, v) = g_{\mu\nu} v^\mu v^\nu$$

Real particles propagate with a **finite 3D speed**, **constant 4D speed**, and the corresponding Euler-Lagrange equations of motion are equivalent to the geodesic equations.

$$S[x(\tau)] = \int_A^B L(x, v) d\lambda$$


$d\tau = L_1(x, v) d\lambda$

$$\left( \frac{\delta S_1[x(\tau)]}{\delta x(\tau)} = 0 \right) \cong \left( \frac{\delta S_2[x(\tau)]}{\delta x(\tau)} = 0 \right) \cong (D_{\vec{v}} \vec{v} = 0)$$

In proper time parameterization  $L=1$  or  $0$        $g_{\mu\nu} v^\mu v^\nu = \text{const}$

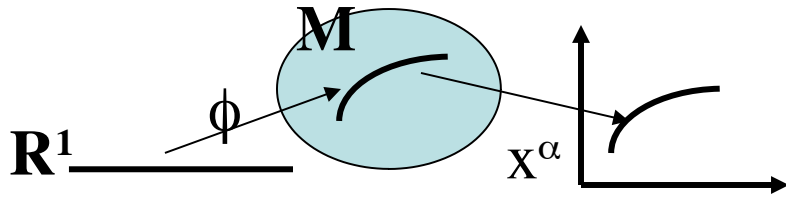
*Should the smallest “distance” be the guiding principle?  
If yes, how it should be defined in other situations?*

# Reparametrization Invariance

Leaving aside expectation that the spacetime is quantized ... we ask:

- Should there be any preferred trajectory parameterization in the 4D spacetime?
- Aren't we free to choose the standard of distance (time)? (Using natural units  $c=1$ )

**Smallest time interval ?!... no more events under certain threshold?**



$$\phi : \mathbb{R}^1 \rightarrow M$$

$$\phi^* : \mathbb{R}^1 \rightarrow TM$$

$$\phi_* : \mathbb{R}^1 \leftarrow T^*M$$

$$\frac{d}{d\tau} \left( \frac{\partial L}{\partial v^\alpha} \right) = \frac{\partial L}{\partial x^\alpha}$$

Integrating using a *Volume* form is natural.

Consider the pull-back of a one-form  $\omega$  :

$$\phi_*(\omega) = A_\alpha(x) v^\alpha d\tau = L(x, v) d\tau,$$

this is first order homogeneous Lagrangian.

If  $L(x, v)$  is homogeneous function of  $v$  then the Euler-Lagrange equations are invariant under  $\tau$ -rescaling,  $L(x, v) \sim H$  is constant.

*What is the general mathematical expression for homogeneous function of first order?*

$$f(\vec{v}) = \sum_i v^i \frac{\partial}{\partial v^i} f(\vec{v})$$

**Is there a unique procedure to extract A, g, S ... from L:**

$$L(x, v) = A_\alpha(x) v^\alpha + \sqrt{g_{\mu\nu}(x) v^\mu v^\nu} + \dots + \sqrt[n]{S(v, v, \dots, v)} + \dots$$

# First Order Homogeneous Lagrangians

## Cons

- Constraints among the Euler-Lagrange equations
  - $(x, v) \rightarrow (x, p)$  problem
  - The Hamiltonian function is ZERO
- Some researchers do not like it...

$$H = v^\alpha p_\alpha - L, p_\alpha = \frac{\partial L}{\partial v^\alpha}$$

## Canonical form:

$$L(x, v) = qA_\alpha(x)v^\alpha + m\sqrt{g_{\mu\nu}(x)v^\mu v^\nu} + \dots + \varepsilon^n \sqrt{S(v, v, \dots, v)} + \dots$$

## Coordinate time

## parameterization $c=1$ :

$$L(x, v) = qA_0(x) + q\vec{A} \cdot \vec{v} + m\sqrt{1 - \vec{v} \cdot \vec{v}} + \dots$$

## Pros

- Reparametrization invariance
- Parameterization independent pathintegral quantization
- May help in dealing with singularities
- Any  $L(x, v) \rightarrow L(x, v/v_0)v_0$  a reparametrization invariant one.
- Easily generalized to p-branes

# Important Lagrangians

➤ 0-brane (point particle):  $\omega^\Gamma \rightarrow \frac{dx^\alpha}{d\tau} = v^\alpha$ ,

$$L(x, \omega) = A_\Gamma(x) \omega^\Gamma + \sqrt{g_{\Gamma\Lambda}(x) \omega^\Gamma \omega^\Lambda} \rightarrow L(x, v) = eA_\alpha(x) v^\alpha + m \sqrt{g_{\mu\nu}(x) v^\mu v^\nu}$$

➤ 1-brane (strings):  $\omega^\Gamma \rightarrow Y^{\alpha\beta} = \frac{\partial(x^\alpha, x^\beta)}{\partial(\tau, \sigma)}$ ,

$$Y^{\alpha\beta} = \det \begin{pmatrix} \partial_\tau x^\alpha & \partial_\sigma x^\alpha \\ \partial_\tau x^\beta & \partial_\sigma x^\beta \end{pmatrix} = \partial_\tau x^\alpha \partial_\sigma x^\beta - \partial_\sigma x^\alpha \partial_\tau x^\beta,$$

**String theory Lagrangian:**  $L(x^\alpha, \partial_i x^\beta) = \sqrt{Y^{\alpha\beta} Y_{\alpha\beta}}$

➤ p-brane: **Dirac-Nambu-Goto Lagrangian:**  $L(x^\alpha, \partial_D x^\beta) = \sqrt{Y^\Gamma Y_\Gamma}$

# One Time Physics, Causality, and $m > 0$

Assume that the gravity like term is always present:  $m \sqrt{g_{\alpha\beta}(x) v^\alpha v^\beta}$

$$g_{\alpha\beta} v^\alpha v^\beta \geq 0$$

*Positivity guarantees well defined proper time! Thus, causality!*

*Select a local coordinate system where  $g$  is  $(+\dots+,-\dots-)$ , call the  $(+)$  coordinates time coordinates and the  $(-)$  space coordinates.*

## Cases:

1. ~~No time coordinates.~~

$$g_{\alpha\beta} v^\alpha v^\beta = - \sum_i (v^i)^2 \leq 0$$

2. Two and more time coordinates.

$$g_{\alpha\beta} v^\alpha v^\beta = (V)^2 + (v^0)^2 - \sum_i (v^i)^2,$$

$$\Rightarrow V^2 + 1 \cong \vec{v}^2$$

3. One time coordinate.

$$g_{\alpha\beta} v^\alpha v^\beta = (v^0)^2 - \sum_i (v^i)^2 \rightarrow 1 - \vec{v}^2 \Rightarrow 1 \geq \vec{v}^2$$

**$m > 0$ , Minkowski signature, and Common Arrow of time! ( $m > 0 \Leftrightarrow |\phi| > 0$  [\*])**

[\*] Reparametrization Invariance and Some of the Key Properties of Physical Systems  
V.G. Gueorguiev and A. Maeder in Symmetry 2021, 13, 522. (10.3390/sym13030522) [ArXiv:1903.02483]

## *Key New Results & Concepts*

- Un-proper time parametrization as source of **extra forces/accelerations**.
- **Pathological behavior** in  $S_n$  interactions for  $n > 2$ .
- Resolving **ambiguities for the interaction fields** using the canonical form of  $L$ .



# *Equivalent Lagrangians*

$$v^i \frac{\partial L(x, \vec{v})}{\partial v^i} = nL(x, \vec{v}), \quad h = pv - L \rightarrow h = (n-1)L$$

$$L \rightarrow f(L), \quad \frac{d}{d\tau} \left( \frac{\partial L}{\partial v} \right) = \frac{\partial L}{\partial x} \rightarrow \frac{d}{d\tau} \left( f' \frac{\partial L}{\partial v} \right) = f' \frac{\partial L}{\partial x}$$

0

$$\frac{f'' \left( \frac{\partial L}{\partial v} \right) \frac{dL}{d\tau}}{f' \left( \frac{\partial L}{\partial v} \right)} + \frac{d}{d\tau} \left( \frac{\partial L}{\partial v} \right) = \frac{\partial L}{\partial x}$$

**Fictitious forces/accelerations in un-proper time parametrization**

$$\frac{dp_\mu}{dt} = \partial_\mu L - \frac{f''}{f'} \dot{L} p_\mu, \quad p_\mu = \frac{\partial L}{\partial v^\mu} \Rightarrow \frac{dp_\mu}{dt} = \partial_\mu L + \frac{\dot{L}}{2L} p_\mu, \quad \text{when } f(L) = \sqrt{L}.$$

$$\frac{dv_\mu}{dt} = \partial_\mu L + \kappa_0 v_\mu, \quad \text{when } L = g_{\mu\nu} v^\mu v^\nu. \quad \text{for } L = \lambda^2 L_{GR} \text{ and } \kappa_0 = -d\lambda/(\lambda dt).$$

# *Pathology of $S_n$ interactions for $n > 2$*

$$L = (S_n(v))^{1/n} \quad \text{parametrization gauge } S_n(v) = \text{const}$$

$$S_n(r, w, u) = \psi(r)w^n + \phi(r)u^n. \quad \text{where } w = dx^0/d\tau \text{ and } u = dr/d\tau$$

$$\frac{du}{d\tau} = -\frac{u^2\phi'(r)}{(n-1)\phi(r)} + \frac{1}{u^{n-2}} \frac{w^n\psi'(r)}{n(n-1)\phi(r)'}$$

$$\frac{dw}{d\tau} = -\frac{wu\psi'(r)}{(n-1)\psi(r)'}$$

For  $n > 2$ , using  $u/w = v/c$  and dividing by  $w^2$ , one see a **pathology at  $v \rightarrow 0$**  (no co-moving rest frame):

$$\frac{(n-1)}{w^2} \frac{du}{d\tau} = -\frac{v^2\phi'(r)}{c^2\phi(r)} + \frac{c^{n-2}}{v^{n-2}} \frac{\psi'(r)}{n\phi(r)'}$$

DM

QM & Inflation

# Summary

- ✓ First order homogeneous *matter Lagrangians* in canonical form naturally contain *electromagnetic and gravitational* interactions.

$$A_\mu(x) = \frac{1}{2}(\tilde{L}(x, v) - \tilde{L}(x, -v))_{/\mu} \Big|_{v=(1, \vec{0})} \quad L = v^\mu A_\mu(x) + \sqrt{g_{\alpha\beta} v^\alpha v^\beta}$$

$$g_{\alpha\beta}(x) = \frac{1}{4}(\tilde{L}(x, v) + \tilde{L}(x, -v))_{/\alpha/\beta}^2 \Big|_{v=(1, \vec{0})}$$

- ✓ The framework predicts **fictitious forces/accelerations** and specific **new interactions** beyond the *electromagnetic and gravitational* interactions!

$$\frac{(n-1)}{w^2} \frac{du}{d\tau} = -\frac{v^2 \phi'(r)}{c^2 \phi(r)} + \frac{c^{n-2}}{v^{n-2}} \frac{\psi'(r)}{n\phi(r)}.$$

$$\frac{dv_\mu}{dt} = \partial_\mu L + \kappa_0 v_\mu, \text{ when } L = g_{\mu\nu} v^\mu v^\nu.$$

# *On the Canonical Form of the First Order Homogeneous Lagrangians*

$$L(x, v) = \frac{h_{\alpha\beta}(x) v^\alpha v^\beta}{\sqrt{g_{\mu\nu}(x) v^\mu v^\nu}},$$

Both fields ( $g$  and  $h$ )  
have same type of  
matter sources  $\sim vv$ .

$$\frac{\delta L(x, v)}{\delta h_{\mu\nu}} = \frac{L(x, v)}{(h_{\alpha\beta}(x) v^\alpha v^\beta)} v^\mu v^\nu, \quad \frac{\delta L(x, v)}{\delta g_{\mu\nu}} = \frac{L(x, v)}{(g_{\alpha\beta}(x) v^\alpha v^\beta)} v^\mu v^\nu.$$

**The canonical form defines fields for each source type  $vvv\dots v$ .**

$$L(x, v) = qA_\alpha(x)v^\alpha + m\sqrt{g_{\mu\nu}(x) v^\mu v^\nu} + \dots + \varepsilon^n \sqrt{S(v, v, \dots, v)} + \dots$$

# *Background Interaction Fields*

*Matter should provide the sources for the interaction fields!*

*Electromagnetic  
interaction*

*Gravitational  
interaction*

*Do these exist?*

$$L(x, v) = qA_{\alpha}(x)v^{\alpha} + m\sqrt{g_{\mu\nu}(x)v^{\mu}v^{\nu}} + \dots + \varepsilon^n \sqrt{S(v, v, \dots, v)} + \dots$$

*Constructing Lagrangians for the interaction fields.*

- ***The “never ending process”:***  
Consider each background field ( $S_n$ ) as M-brane theory since  $S_n : M \rightarrow \mathcal{S}^n M$ .
- ***External algebra method:*** (*A-type fields*)  
Constructing external n-forms proportional to the volume form using external derivatives and the Hodge dual operation:  $A \wedge *A$  and  $dA \wedge *dA$ .
- ***Gauge symmetry method:***  
*Use the Euler-Lagrange equations or their integral (geodesic) deviation.*

# Field Lagrangians

**Lorentz equation:**  $\vec{a} = \hat{F} \cdot \vec{v}, \hat{F} = dA + [A, A]$

**Geodesic equation:**  $\vec{a} = \vec{v} \cdot \Gamma \cdot \vec{v}, \Gamma \rightarrow \Gamma + \partial\Lambda$

**Geodesic deviation:**  $\frac{d}{d\tau} \frac{d}{d\tau} \vec{\xi} = \hat{R} \vec{\xi}, \hat{R} = d\Gamma + [\Gamma, \Gamma]$

$A \wedge *A$  does not obey the symmetries of the matter equations,

when  $F \wedge F = d(A \wedge F)$  is a boundary term... (Q-Hall effect)  $L(A, \partial A) = \int_M F \wedge *F$

$R$  is a **special** curvature two-form with values in  $\text{Lie}(TM)$ .

$$R_{ij,kl} = -R_{ij,lk} = -R_{lk,ij} = R_{kl,ij}, \quad R^* = R_{ij,*(jk)}, \quad \boxed{R^\wedge = R_{i[j,k]l}} \quad ? \quad (4D)$$

**Einstein-Cartan action:**  $L(g, \partial g) = \int_M \hat{R}_{\alpha\beta} \wedge * (dx^\alpha \wedge dx^\beta)$

# Summary

- ✓ First order homogeneous *matter Lagrangians* in *canonical form* naturally contain *electromagnetic and gravitational* interactions.
- ✓ Gravitational term with **Minkowski signature** and **positive mass** is needed to assure **well-defined proper time** and **causality**; thus, **common arrow of time**.
- ✓ Unique construction of *Lagrangians for the interaction fields*.
- ✓ The framework predicts specific **new interactions** beyond the *electromagnetic and gravitational* interactions!