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Centerless BMS4 charge algebra and (A)ds uplift

Geoffrey Compère Université Libre de Bruxelles (ULB)

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1. The centerless BMS4 charge algebra

Infrared structure of gravity



"The Hamiltonian in General Relativity is a surface term. Therefore, gravity is holographic."

Infrared structure of gravity



AdS case $\Lambda < 0$.

Flat case $\Lambda = 0$.

60'S [ADM ; BMS]

dS case $\Lambda > 0$.

2000's

80's

[Ashtekar,Brown,Bunster,Henneaux] [Maldacena,Witten]

2010-2021 [Barnich-Troessaert ; Strominger et al]

Boundary conditions -> Global symmetries



Dirichlet Anti-de Silter

SOLUTION

SYMMETRY

N

CHARGES

(1)



AdS case $\Lambda < 0$.

$$ds^{2} = -\frac{3}{\Lambda} \frac{d\rho^{2}}{\rho^{2}} + \gamma_{ab}(\rho, x^{c}) dx^{a} dx^{b}.$$
[Fefferman-Graham theorem]
$$\gamma_{ab} = \frac{1}{\Lambda} g_{ab}^{(0)} + \frac{1}{\Lambda} g_{ab}^{(1)} + g_{ab}^{(2)} + \rho g_{ab}^{(3)} + \mathcal{O}(\rho^{2})$$

$$T = \sqrt{3|\Lambda|} (3)$$

 $I_{ab} = \frac{1}{16\pi G} g_{ab}$

 $D_a^{(0)}T^{ab} = 0, \quad g_{ab}^{(0)}T^{ab} = 0.$

Two holographic fields: the boundary metric $g_{ab}^{\left(0
ight)}$ and the stress-tensor T^{ab}

Fixing the boundary metric to be the flat cylinder, there are SO(2,3) symmetries.

 $\mathcal{L}_{\xi^{(0)}} g_{ab}^{(0)} \sim g_{ab}^{(0)} \qquad \mathcal{L}_{\xi} g_{\mu\nu} = g_{\mu\nu} (T_{ab} + \delta_{\xi} T_{ab}) - g_{\mu\nu} (T_{ab}).$

The associated charges are conserved and represent the group SO(2,3) under the Peierls bracket

 $Q_{\xi} = \int_{S} d^{2}\Omega T_{ab} \xi^{a}_{(0)} n^{b}, \qquad \{Q_{\xi}, Q_{\eta}\} = Q_{[\xi, \eta]}$

Asymptotically Flat Spacetimes: Null infinity



Flat case $\Lambda = 0$.

Gauge fixing : Bondi / Newman-Unti coordinates $g_{ur} = -1, \quad g_{uu} = 0, \quad g_{uA} = 0, \quad x^A = \{\theta, \phi\}$

$$ds^{2} = e^{2\beta} \frac{V}{r} du^{2} - 2e^{2\beta} du dr + g_{AB} (dx^{A} - U^{A} du) (dx^{B} - U^{B} du)$$

$$g_{AB} = r^2 q_{AB} + r C_{AB} + D_{AB} + \frac{1}{r} E_{AB} + \frac{1}{r^2} F_{AB} + \mathcal{O}(r^{-3})$$

Infinite number of holographic fields: the boundary metric q_{AB} , the shear $C_{AB}(u, x^C)$, mass $m(u, x^C)$ and angular momentum aspects $N_A(u, x^C)$, subleading fields $E_{AB}(u, x^C)$, $F_{AB}(u, x^C)$

Flux-balance laws:

NOTLONOS

4

 $\partial_u m + D_A(\cdots) = \text{HARD TERMS}(N_{AB}),$ $\partial_u N_A + D_B(\cdots) = \text{HARD TERMS}(N_{AB}),$



SYMMETRY

N

I. Supertranslations

$$T(\theta,\phi)\partial_u + \frac{1}{2}\nabla^2 T \partial_r - \frac{1}{r}(\partial_\theta T \partial_\theta + \frac{1}{\sin^2 \theta}\partial_\phi T \partial_\phi) + \dots$$

- The associated Noether charge is the Bondi mass aspect m(u,x^A) integrated over the celestial sphere
- The 4 lowest harmonics are the translations associated with Momenta.
- Supertranslations transitions are associated with displacement memory and are caused by any null radiation exiting null infinity

II. Super-Lorentz transformations

$$\frac{1}{2}uD_AR^A\partial_u + \left(-\frac{1}{2}(r+u)D_AR^A + \mathcal{O}(\frac{1}{r})\partial_r + \left(R^A - \frac{u}{2r}D^AD_BR^B + \mathcal{O}(\frac{1}{r^2})\right)\partial_A$$

- Associated Noether charge: NA(u,x^A) integrated over the celestial sphere (after renormalization of radial divergences)
- The 6 lowest harmonics are associated with the Lorentz charges: angular momentum and center-of-mass charge (orbital angular momentum).
- Lorentz transformations are asymptotic symmetries. Super-Lorentz transformations are asymptotic symmetries after renormalization.
- Superrotations and superboosts

Symmetry group = $Vect(S^2) \ltimes Diff(S^2)$ Generalized BMS4 group



3. CHARGES



 $Q_T(u) = \int_S d^2 S \,\overline{m}(u, x^C) T(x^C)$ $Q_R(u) = \int_S d^2 S \,\overline{N}_A(u, x^C) R^A(x^C)$

Junction condition between past and future null infinity: Antipodal map at spatial infinity

Scattering around Minkowski obeys $F_{T,R} = \int_{-\infty}^{\infty} du \partial_u Q_{T,R}^+(u) = \int_{-\infty}^{\infty} dv \partial_v Q_{T,R}^-(v)$

This is the Ward identity of BMS symmetry. It is equivalent to the leading and subleading soft theorems.

BMS4 flux asymptotic symmetry algebra



Flat case $\Lambda = 0$.

Prescription for the experts:

 $\bar{M} = M + \frac{1}{8}C_{AB}N_{\text{vac}}^{AB},$ $\bar{N}_A = N_A^{[B.T]} u \partial_A \bar{M} + \frac{1}{4}C_{AB}D_C C^{BC} + \frac{3}{32}\partial_A (C_{BC}C^{BC}).$ Given a prescription

CHARGE

cŋ

 $\overline{m} = m + f(q_{AB}, C_{AB}, N_{AB}),$ $\overline{N}_A = N_A + f_A(q_{AB}, C_{AB}, N_{AB})$

and boundary conditions at past/future times.

The BMS4 algebra can be represented under the Peierls bracket without central extension

 $\{F_{T_1}, F_{T_2}\} = 0,$

 $\{F_{R_1}, F_{T_2}\} = F_{R_1(T_2)}, \qquad R_1(T_2) \equiv (R_1^A \partial_A - \frac{1}{2} D_A R_1^A) T_2$

 $\{F_{R_1}, F_{R_2}\} = F_{[R_1, R_2]}$

[Campiglia, Peraza, 2020] [G.C., Fiorucci, Ruzziconi, 2020]

Its quantization leads to the soft graviton theorems.

2. The angular momentum in CrR

Three ambiguities to define J

1. Center-of-mass frame

 $SO(2) \subset SO(3) \subset \overline{SO(3,1)}$

- Pauli-Lubanski spin pseudo-vector

$$W_{\mu} = \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} J^{\nu\rho} P^{\sigma}$$

 $k^{\mu} = (1, n_i)$

- Local rotation vector :

$$R_i^{\prime A} = \gamma R_i^A + (1-\gamma) \frac{v_i (\vec{v} \cdot \vec{R}^A)}{v^2} + \gamma \epsilon_{ijk} v_j K_k^A, \qquad v_i \equiv \frac{\mathcal{P}_i}{\mathcal{P}_0}, \qquad \gamma = (1 - \frac{v^2}{c^2})^{-\frac{1}{2}}, \quad v = \sqrt{\vec{v} \cdot \vec{v}}.$$

In GR, the 4-momentum evolves according to the mass loss formula:

$$\dot{\mathcal{P}}^{\mu} = -\frac{c^2}{8G} \oint_S \dot{C}_{AB} \dot{C}^{AB} k^{\mu}$$

Three ambiguities to define J 2. Supertranslation frame $SO(2) \subset SO(3) \subset SO(3,1) \subset SO(3,1) \rtimes Vect(S^2)$

- Boundary condition on the shear $C_{AB}|_{u=\pm\infty} = -2D_A D_B C^{\pm} + \gamma_{AB} D^C D_C C^{\pm} + O(u^{-1}).$ $\delta C^{\pm} = T(\theta, \phi)$

- Fix supertranslation frame at $r o \infty, u o -\infty (\mathcal{I}_{-}^{+})$



- The displacement memory effect is generally present

Three ambiguities to define J3. α -ambiguity

$$\mathcal{J}_{i}^{(\alpha)} \equiv -\frac{1}{2} \oint_{S} \epsilon^{AD} \partial_{D} n_{i} \left(\bar{N}_{A} - \frac{\alpha c^{3}}{4G} C_{AB} D_{C} C^{BC} \right)$$

For all α

(i) Vanishing for Minkowski
(ii) Standard J of Kerr
(iii) Locally constructed from tensors
(iv) Obey the BMS algebra
(v) Satisfy the BMS flux-balance laws

Three ambiguities to define J3. α -ambiguity

$$\mathcal{J}_{i}^{(\alpha)} \equiv -\frac{1}{2} \oint_{S} \epsilon^{AD} \partial_{D} n_{i} \left(\bar{N}_{A} - \frac{\alpha c^{3}}{4G} C_{AB} D_{C} C^{BC} \right)$$

Definitions used in the literature:

$$\alpha = 1 \qquad \mathcal{J}_i = \frac{1}{16\pi G} \oint_S D^{\mu} \xi_i^{\nu} \frac{1}{2} \epsilon_{\mu\nu\alpha\beta} dx^{\alpha} \wedge dx^{\beta}$$

[Komar][Iyer,Wald,1992] [Wald,Zoupas,1999]

$$\dot{\mathcal{J}}_{i} = -\frac{G}{c^{5}} \left(\frac{2}{5} \epsilon_{ijk} I_{jl}^{(2)} I_{kl}^{(3)} \right) - \frac{G}{c^{7}} \left(\frac{1}{63} \epsilon_{ijk} I_{jlm}^{(3)} I_{klm}^{(4)} + \frac{32}{45} \epsilon_{ijk} J_{jl}^{(2)} J_{kl}^{(3)} \right) + O\left(c^{-9}\right), \quad \begin{bmatrix} \text{Thorne} I \\ 1980 \end{bmatrix}$$

$$\dot{\mathcal{J}}_i = \frac{c^3}{32\pi G} \oint_S d^2 \Omega \epsilon_{ijk} (x^i \dot{f}_{ab} \partial_j f_{ab} - 2f_{ia} \dot{f}_{ja})$$

$$\dot{\mathcal{J}}_i = \frac{1}{16\pi G} \oint_S d^2 \Omega(\mathcal{L}_{\xi_i} D_c - D_c \mathcal{L}_{\xi_i}) l_d q^{ac} q^{bd}$$

[Landau-Lifshitz]

[Dray-Streubel,84] [AshEekar, Streubel,81]

(vii) No background structure required (vii) Axisymmetry implies J=0 Three ambiguities to define J3. α -ambiguity

$$\mathcal{J}_{i}^{(\alpha)} \equiv -\frac{1}{2} \oint_{S} \epsilon^{AD} \partial_{D} n_{i} \left(\bar{N}_{A} - \frac{\alpha c^{3}}{4G} C_{AB} D_{C} C^{BC} \right)$$

Definitions used in the literature:

α = 0 [Strominger, Zhiboedov, 2014] [Pasterski, Strominger, Zhiboedov, 2015] [G.C., Fiorucci, Ruzziconi, 2020]

The change of definition leads to a numerically 0.01%-0.1% effect for binary coalescences [Elhashash, Nichols, 2021]

(vi) Background structure required (radial foliation) (vii) Generalized BMS group represented (including super-Lorentz) $\{F_{T_1}, F_{T_2}\} = 0, \quad \{F_{R_1}, F_{T_2}\} = F_{R_1(T_2)}, \qquad \{F_{R_1}, F_{R_2}\} = F_{[R_1, R_2]}$ $R_1(T_2) \equiv (R_1^A \partial_A - \frac{1}{2}D_A R_1^A)T_2$

3. Extension of the BMS group to (A)ds



• Three-dimensional case :



• Four-dimensional case :



Universal BMS structure (keeping all dynamics)



A dictionary exists between distinct bulk gauges

Definitions

Starobinsky/ Fefferman-Graham (SFG) gauge

 (ρ, x^a)

$$g_{\rho a} = 0,$$
$$g_{\rho \rho} = -\frac{3}{\Lambda} \frac{1}{\rho^2}$$

Bondi gauge

 (u, r, x^A) $g_{rr} = 0, \quad g_{rA} = 0$

 $\partial_r \left(\frac{\det(g_{AB})}{r^4} \right) = 0$

The dictionary between Bondi and Starobinsky/Fefferman-Graham gauge has been worked out

- One can solve the large radius
 expansion of Einstein's equations in
 both gauges
- @ A diffeomorphism exists between the two gauges when $\Lambda \neq 0$
- The (2-covariant) map between the free fields in each gauge can be formulated

[Poole, Skenderis, Taylor, 2018] [G.C., Fiorucci, Ruzziconi, 2019]

Solution space (Al(A)dS4)
SFG gauge

$$ds^{2} = -\frac{3}{\Lambda}\frac{d\rho^{2}}{\rho^{2}} + \gamma_{ab}(\rho, x^{c})dx^{a}dx^{b}.$$

 $ds^{2} = e^{2\beta}\frac{V}{r}du^{2} - 2e^{2\beta}dudr + g_{AB}(dx^{A} - U^{A}du)(dx^{B} - U^{B}du)$
 $g_{AB} = r^{2}q_{AB} + rC_{AB} + D_{AB} + \frac{1}{r}E_{AB} + \frac{1}{r^{2}}F_{AB} + \Theta(r^{-3})$

$$\gamma_{ab} = \frac{1}{\rho^2} g_{ab}^{(0)} + \frac{1}{\rho} g_{ab}^{(1)} + g_{ab}^{(2)} + \rho g_{ab}^{(3)} + \mathcal{O}(\rho^2)$$

[FG theorem]

$$T_{ab} = \frac{\sqrt{3|\Lambda|}}{16\pi G} g_{ab}^{(3)}$$

$$l = l_A^A = \frac{1}{2} q^{AB} \partial_u q_{AB} = \partial_u \ln \sqrt{q}.$$

$$U^{A} = U_{0}^{A}(u, x^{B}) + U^{(1)}(u, x^{B})\frac{1}{r} + U^{(2)}(u, x^{B})\frac{1}{r^{2}} + U^{(3)}(u, x^{B})\frac{1}{r^{3}} + U^{(L3)}(u, x^{B})\frac{\ln r}{r^{3}} + o(r^{-3})$$

$$\beta(u, r, x^{A}) = \beta_{0}(u, x^{A}) + \frac{1}{r^{2}} \Big[-\frac{1}{32} C^{AB} C_{AB} \Big] + \frac{1}{r^{3}} \Big[-\frac{1}{12} C^{AB} \mathcal{D}_{AB} \Big] + \frac{1}{r^{4}} \Big[-\frac{3}{32} C^{AB} \mathcal{E}_{AB} - \frac{1}{16} \mathcal{D}^{AB} \mathcal{D}_{AB} + \frac{1}{128} (C^{AB} C_{AB})^{2} \Big] + \mathcal{O}(r^{-5}).$$

$$\begin{aligned} \frac{V}{r} &= \frac{\Lambda}{3} e^{2\beta_0} r^2 - r(l + D_A U_0^A) \\ &- e^{2\beta_0} \Big[\frac{1}{2} \Big(R[q] + \frac{\Lambda}{8} C_{AB} C^{AB} \Big) + 2D_A \partial^A \beta_0 + 4\partial_A \beta_0 \partial^A \beta_0 \Big] - \frac{2M}{r} + o(r^{-1}) \\ \frac{\Lambda}{3} C_{AB} &= e^{-2\beta_0} \Big[(\partial_u - l) q_{AB} + 2D_{(A} U_{B)}^0 - D^C U_C^0 q_{AB} \Big]. \end{aligned}$$

Holographic fields
$$\Lambda \neq 0$$

SFG gauge
 $g_{ab}^{(0)} dx^a dx^b$
 T^{ab}
 $g_{ab}^{(0)} T^{ab} = 0.$
Bondi gauge
 $(2+1 \text{ boundary split})$
 $\frac{\Lambda}{3}e^{4\beta_0}du^2 + q_{AB}(dx^A - U_0^A du)(dx^B - U_0^B du)$
 $T_{ab} = \frac{\sqrt{3}[\Lambda]}{16\pi G} \begin{bmatrix} -\frac{4}{3}M^{(\Lambda)} & -\frac{2}{3}N_B^{(\Lambda)} \\ -\frac{2}{3}N_A^{(\Lambda)} & J_{AB} + \frac{2}{3}M^{(\Lambda)}q_{AB} \end{bmatrix}$
Boundary ODES

$$\begin{split} N_A^{(CV)} &= N_A - \frac{1}{2\Lambda} D^D (N_{AB} - \frac{1}{2} l C_{AB}) - \frac{1}{4} \partial_A (\frac{1}{\Lambda} R[q] - \frac{1}{8} C_{CD} C_{AB} \\ J_{AB} &= - \mathcal{E}_{AB} - \frac{3}{\Lambda^2} \Big[\partial_u (N_{AB} - \frac{1}{2} l C_{AB}) - \frac{\Lambda}{2} q_{AB} C^{CD} (N_{CD} - \frac{1}{2} l C_{AB} \\ &+ \frac{3}{\Lambda^2} (D_A D_B l - \frac{1}{2} q_{AB} D_C D^C l) \\ &- \frac{1}{\Lambda} (D_{(A} D^C C_{B)C} - \frac{1}{2} q_{AB} D^C D^D C_{CD}) \\ &+ C_{AB} \Big[\frac{5}{16} C_{CD} C^{CD} + \frac{1}{2\Lambda} R[q] \Big]. \end{split}$$

$$D_a^{(0)}T^{ab} = 0$$

$$(\partial_u + \frac{3}{2}l)M^{(\Lambda)} + \frac{\Lambda}{6}D^A N^{(\Lambda)}_A + \frac{\Lambda^2}{24}C_{AB}J^{AB} = 0.$$
$$(\partial_u + l)N^{(\Lambda)}_A - \partial_A M^{(\Lambda)} - \frac{\Lambda}{2}D^B J_{AB} = 0.$$

More holographic fields for $\Lambda=0$ sfe gauge Bondi gauge $\frac{\Lambda}{3}e^{4\beta_0}du^2 + q_{AB}(dx^A - U_0^A du)(dx^B - U_0^B du)$ $M, N_A, C_{AB}, D_{AB}, E_{AB}, F_{AB}, \cdots$ Constraints $\partial_u q_{AB} = lq_{AB} + 2D_{(A}U^0_{B)} - D^C U^0_C q_{AB}$ $\partial_u M = -\frac{1}{8}N_{AB}N^{AB} + \frac{1}{4}D_A D_B N^{AB} + \frac{1}{8}D_A D^A \mathring{R},$ $\partial_u N_A = D_A M + \frac{1}{16} D_A (N_{BC} C^{BC}) - \frac{1}{4} N^{BC} D_A C_{BC}$ $-\frac{1}{4}D_B(C^{BC}N_{AC} - N^{BC}C_{AC}) - \frac{1}{4}D_BD^BD^CC_{AC}$ $+\frac{1}{4}D_B D_A D_C C^{BC} + \frac{1}{4}C_{AB}D^B \mathring{R}.$ infinite number of $\partial_u D_{AB} = 0,$ boundary ODEs / $\partial_u E_{AB} = \cdots$ flux-balance laws $\partial_u F_{AB} = \cdots$ see [Barnich, Troessart, 2012]

Boundary gauge condition: Definition of Λ -BMS

$$g_{tt}^{(0)} = \frac{\Lambda}{3}, \qquad g_{tA}^{(0)} = 0, \qquad \det(g_{(0)}) = \frac{\Lambda}{3}\bar{q}.$$

$$ds_{(0)}^2 = \frac{\Lambda}{3}dt^2 + q_{AB}dx^A dx^B$$

o Can always be reached

@ Does not constraint the Cauchy problem

The residual diffeomorphisms in a given bulk gauge and in this boundary gauge form the A-BMS group.

Symmetry generators

Preserving SFG gauge: Diff(53) x Weyl

$$\begin{split} \xi^{u} &= f, \\ \xi^{A} &= Y^{A} + I^{A}, \quad I^{A} = -\partial_{B}f \int_{r}^{\infty} dr' (e^{2\beta}g^{AB}), \\ \xi^{r} &= -\frac{r}{2} (\mathcal{D}_{A}Y^{A} - 2\omega + \mathcal{D}_{A}I^{A} - \partial_{B}fU^{B} + \frac{1}{2}fg^{-1}\partial_{u}g) \end{split}$$

 $\partial_r f = 0 = \partial_r Y^A$

Preserving further boundary gauge:

 $g = \det(g_{AB})$

Flat spacetime limit: $Y^A = V^A(x^B), f = T(x^A) + \frac{u}{2}D_A V^A$

(Soft) Algebra / Algebroid

$$\bar{\xi} = f\partial_u + Y^A \partial_A$$

$$[\bar{\xi}_1, \bar{\xi}_2] = \hat{\bar{\xi}},$$

$$\hat{\bar{\xi}} = \hat{f}\partial_u + \hat{Y}^A \partial_A$$

$$\begin{split} \hat{f} &= Y_1^A \partial_A f_2 + \frac{1}{2} f_1 D_A Y_2^A - (1 \leftrightarrow 2), \\ \hat{Y}^A &= Y_1^B \partial_B Y_2^A - \frac{\Lambda}{3} f_1 q^{AB} \partial_B f_2 - (1 \leftrightarrow 2). \end{split}$$

The structure constants of the $\Lambda\text{-BMS}$ algebra are field-dependent.

In the flat limit, the structure constants are field-independent and reproduce the generalized BMS algebra.

A-BMS: Surface charges

$$\mathrm{d}\boldsymbol{k}_{\xi,\mathrm{ren}}[\delta\phi;\phi] = \boldsymbol{\omega}_{\mathrm{ren}}[\delta_{\xi}\phi,\delta\phi;\phi],$$

$$\delta H_{\xi}[\phi] = \int_{S_{\infty}^{2}} 2(\mathrm{d}^{2}x)_{\rho t} \left[\delta \left(\sqrt{|g^{(0)}|} T_{(\mathrm{tot})b}^{t} \right) \xi_{(0)}^{b} - \frac{1}{2} \sqrt{|g^{(0)}|} \xi_{(0)}^{t} T_{(\mathrm{tot})}^{bc} \delta g_{bc}^{(0)} \right].$$

Charges are finite thanks to renormalization

 \circ Charges are neither conserved or integrable $\delta H_{\xi}[\phi] = \delta H_{\xi}[\phi] + \Xi_{\xi}[\delta\phi;\phi],$

Charges associated with Weyl are zero

They obey the surface charge algebra

 $\{H_{\xi}[\phi], H_{\chi}[\phi]\}_{\star} = H_{[\xi, \chi]_{\star}}[\phi].$

where the Barnich-Troessaert brackets are

 $[\xi, \chi]_{\star} = [\xi, \chi] - \delta_{\xi} \chi + \delta_{\chi} \xi.$

 $\{H_{\xi}[\phi], H_{\chi}[\phi]\}_{\star} \equiv \delta_{\chi} H_{\xi}[\phi] + \Xi_{\chi}[\delta_{\xi}\phi, \phi].$

Completeness of Λ -BMS Number of flux-balance laws: d-1

$$(\partial_u + \frac{3}{2}l)M^{(\Lambda)} + \frac{\Lambda}{6}D^A N_A^{(\Lambda)} + \frac{\Lambda^2}{24}C_{AB}J^{AB} = 0.$$
$$(\partial_u + l)N_A^{(\Lambda)} - \partial_A M^{(\Lambda)} - \frac{\Lambda}{2}D^B J_{AB} = 0.$$

Number of generators: d-1 : f, YA Number of charges: d-1 : $M^{(\Lambda)}$, $N^{(\Lambda)}_A$

In comparison with $Diff(S^3)$ studied in [Anninos,Ng,Strominger,2011] the Λ -BMS groupoid is the subset of $Diff(S^3)$ associated with non-trivial flux-balance laws / Ward identities

(A)ds equivalent of Bondi shear/news

Action with holographic counterterms

$$S = \frac{1}{16\pi G} \int_{\mathscr{M}} d^4x \sqrt{-g} \left(R[g] - 2\Lambda \right) + \frac{1}{16\pi G} \int_{\mathscr{I}} d^3x \sqrt{|\gamma|} \left(2K + \frac{4}{\ell} - \ell R[\gamma] \right).$$
 [Balasubramanian, Kraus]

[G.C., Marolf, 2008]

At the boundary, the orthogonal component of the symplectic structure is

$$\omega^{\rho} = \frac{1}{2\ell^2} \int_{\mathscr{I}} d^3x \,\delta\left(\sqrt{|g_{(0)}|}T^{ab}\right) \wedge \delta g_{ab}^{(0)}.$$

$$T_{ab} = \frac{\sqrt{3|\Lambda|}}{16\pi G} \begin{bmatrix} -\frac{4}{3}M^{(\Lambda)} & -\frac{2}{3}N_B^{(\Lambda)} \\ -\frac{2}{3}N_A^{(\Lambda)} & J_{AB} + \frac{2}{\Lambda}M^{(\Lambda)}q_{AB} \end{bmatrix}$$

$$\omega^{\rho} = \frac{3}{32\pi G\ell^4} \int_{\mathscr{I}} d^3x \,\sqrt{\bar{q}} \,\delta J^{AB} \wedge \delta q_{AB}.$$

Therefore, energy is transferred due to changes of both J^{AB} and q_{AB}. They are the analogue in dS/AdS of the Bondi shear/news. [G.C., Fiorucci, Ruzziconi, 2019]

symplectic flux and the Cauchy problem

 $\mathsf{AdS} \ \mathsf{case} \ \Lambda < \mathsf{0}. \qquad \qquad \mathsf{Flat} \ \mathsf{case} \ \Lambda = \mathsf{0}. \qquad \qquad \mathsf{dS} \ \mathsf{case} \ \Lambda > \mathsf{0}.$

In AdS, the Cauchy problem requires an additional boundary condition (standard or "leaky")

Example of "leaky" boundary condition

Fix the initial data of both Ads's. Then the Cauchy problem of the first Ads is well-defined.

Related example: [Almheiri, Hartman, Maldacena, Shaghoulian, Tajdini, 2019]

Conclusion

- The extended BMS charge algebra (supertranslations and Diff(S²) super-Lorentz transformations) is realized without center at the past and future of null infinity (at spatial and timelike infinity). The asymptotic symmetry group at spatial infinity therefore includes the extended BMS group.
- The extended BMS asymptotic symmetry algebra leads to a preferred definition of the quantized angular momentum, which differs from many existing classical prescriptions.
- The extended BMS charge algebra admits a natural extension to (A)dS: the Λ -BMS algebroid. It is the asymptotic symmetry group of Al(A)dS spacetimes with "leaky boundary conditions": without intrinsic boundary conditions (except a boundary gauge fixing condition that does not constraint the Cauchy problem) but with external boundary conditions.